Studying plasmas with computers

Jannis Teunissen Multiscale Dynamics



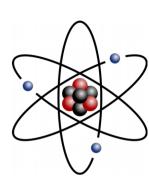
What is a plasma?

- Solids, liquids & gases: electrons stick to atoms
- With enough energy, a **plasma** forms: A gas of ions, electrons (and neutral particles)

- Free charges \rightarrow electric fields
- Moving charges → *magnetic fields*
- Lorentz force

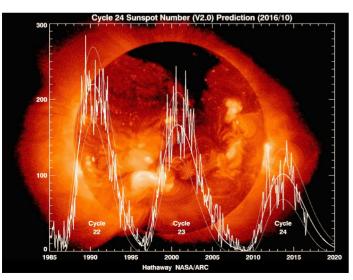
 $\mathbf{F} = q(\mathbf{E} + \mathbf{v} \times \mathbf{B})$





Our plasma energy source: Sun

- Open questions:
 - Coronal heating
 - Solar cycle

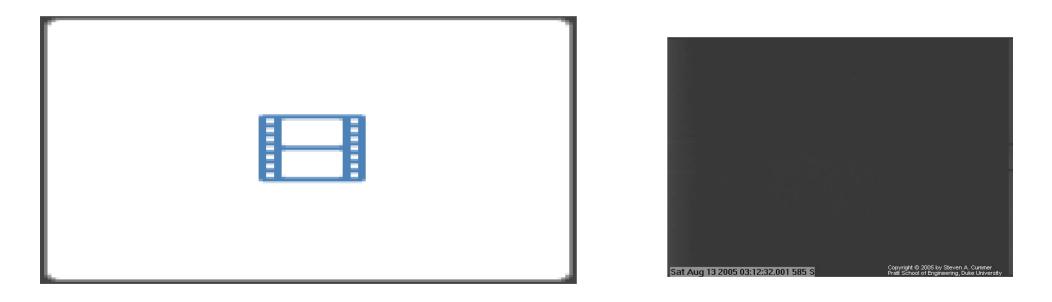


Coronal Mass Ejection





Plasmas in the sky Lightning Sprites



[Wu et al., GRL, 2019]

Technological plasmas

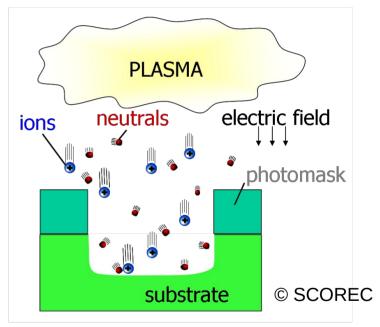
Fluorescent lamps in M133



High-voltage technology



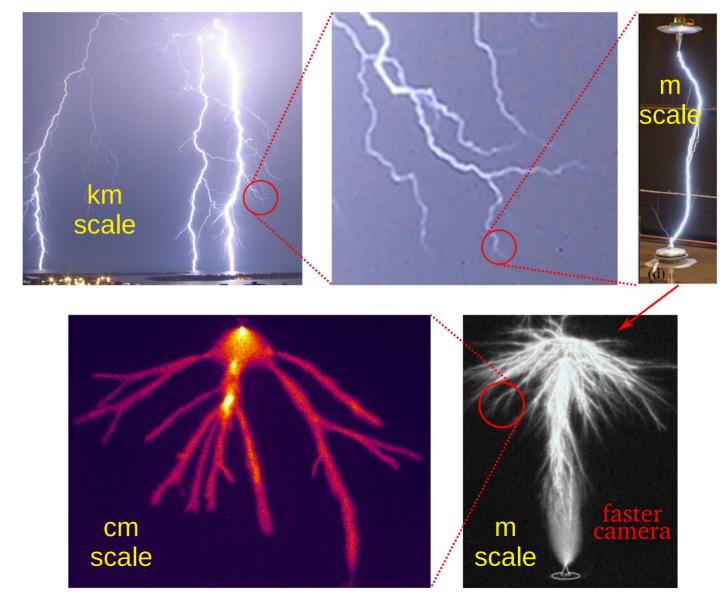
Plasma etching



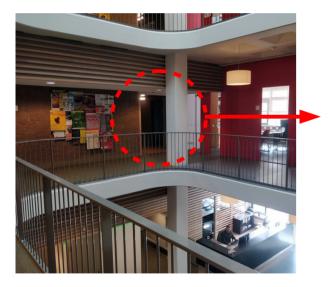




Plasmas are often *multiscale*



Multiscale Dynamics group



Researchers

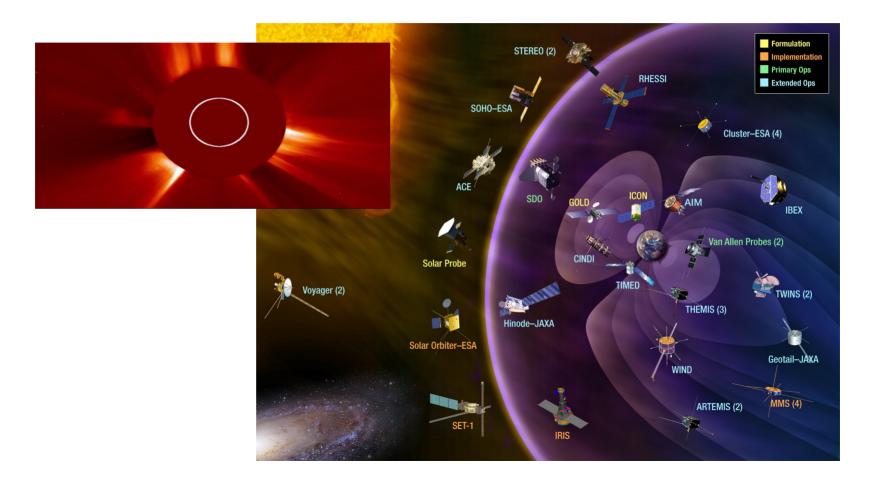
Ute Ebert (leader) Behnaz Bagheri (pd) Hani Francisco (phd) Andy Martinez (phd) Hemaditya Malla (phd) Dennis Bouwman (phd) Baohong Guo (phd) Yaogong Wang (visiting) Jannis Teunissen (staff) Andong Hu (pd) Rakesh Sarma (pd) Carl Shneider (pd) Mandar Chandorkar (phd) Enrico Camporeale (staff, 25%) electric discharges (modeling)



space weather (machine learning)



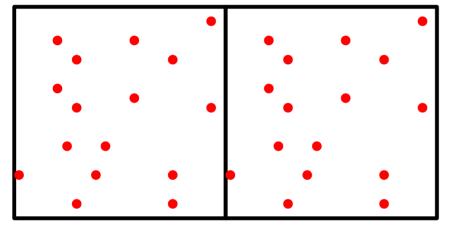
Machine learning for space weather

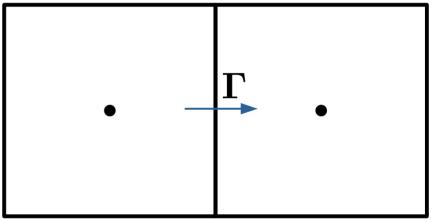


Physical modeling

Particle models





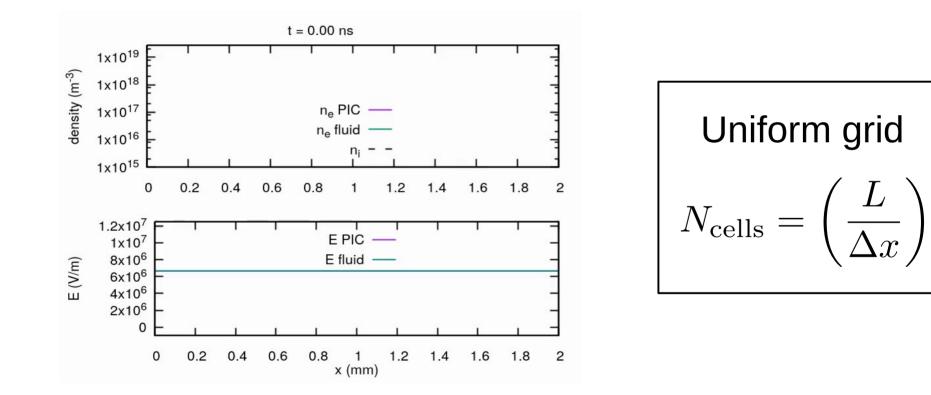


$$\partial_t \mathbf{x} = \mathbf{v}$$

 $\partial_t \mathbf{v} = \mathbf{a}$

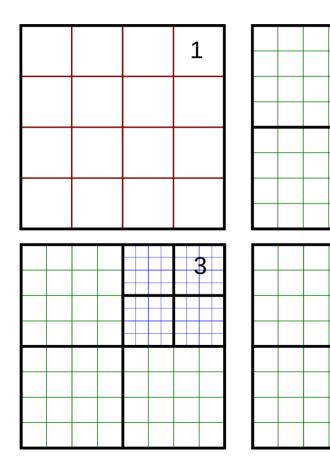
$$\partial_t \mathbf{u} + \nabla \cdot \mathbf{\Gamma} = S$$

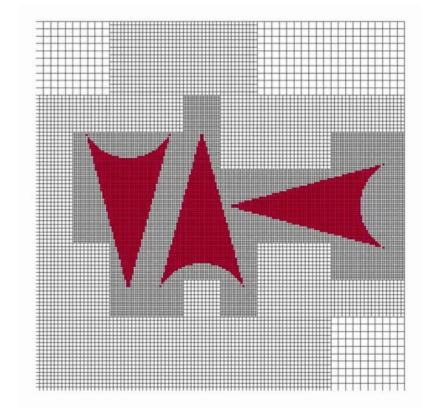
Plasmas often have steep gradients



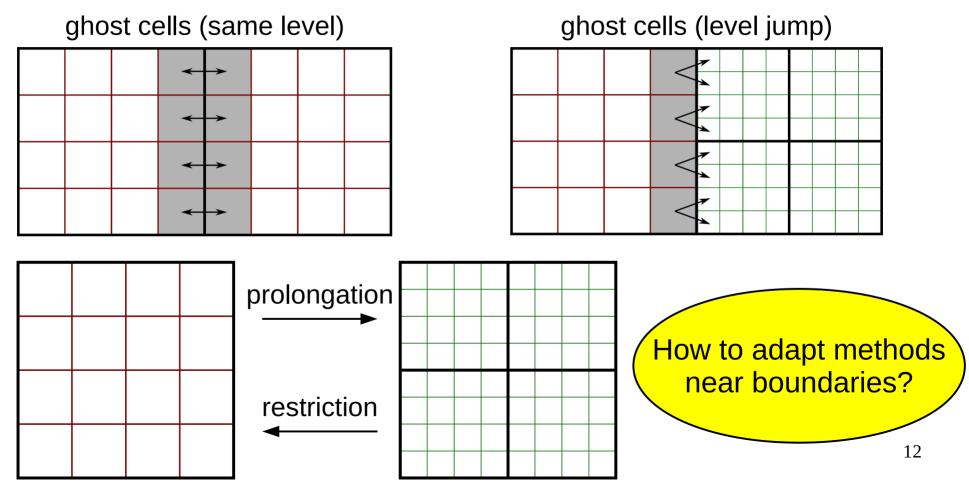
Solution: Adaptive Mesh Refinement

 \mathbf{T}





Parallel communication



Elliptic Partial Differential Equations

Electrostatics

$$\nabla \cdot (\varepsilon \nabla \phi) = -\rho$$

Gravity $abla^2\phi=4\pi G
ho$

Diffusion (particles, heat, radiation)

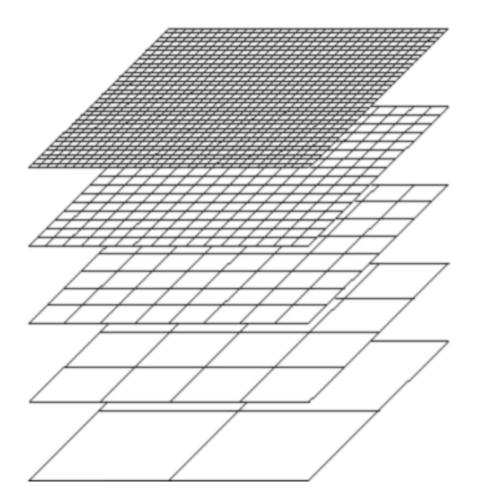
$$\partial_t n = \nabla \cdot (D \nabla n)$$

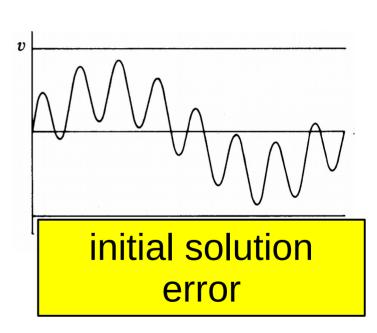
Divergence cleaning Incompressible flow $\partial_t \rho + \nabla \cdot (\rho \mathbf{v}) = 0$ $\nabla \cdot \mathbf{v} = 0$ Magnetohydrodynamics (MHD) $\nabla \cdot \mathbf{B} = 0$

We need *fast* elliptic solvers

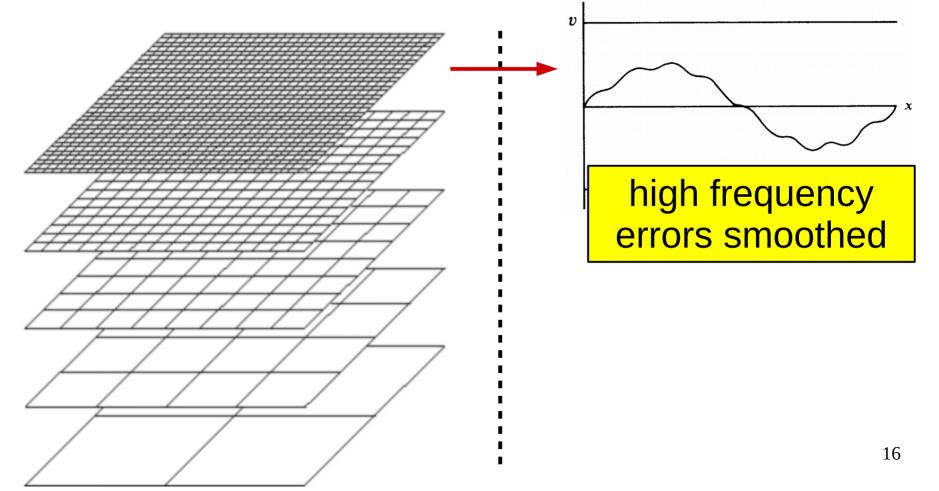
- Direct sparse solvers (e.g. MUMPS, Pardiso)
 - Too expensive in 3D
- FFT-based solvers (e.g. P3DFFT, AccFFT)
 - O(N log N), uniform grid only
- Geometric multigrid (e.g. Hypre, HPGMG)
 - O(N) scaling, AMR compatible

Geometric multigrid

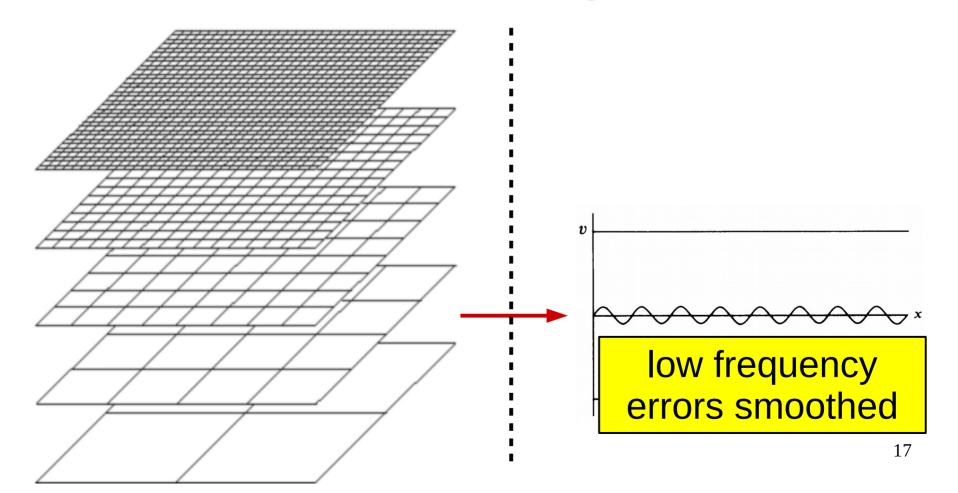




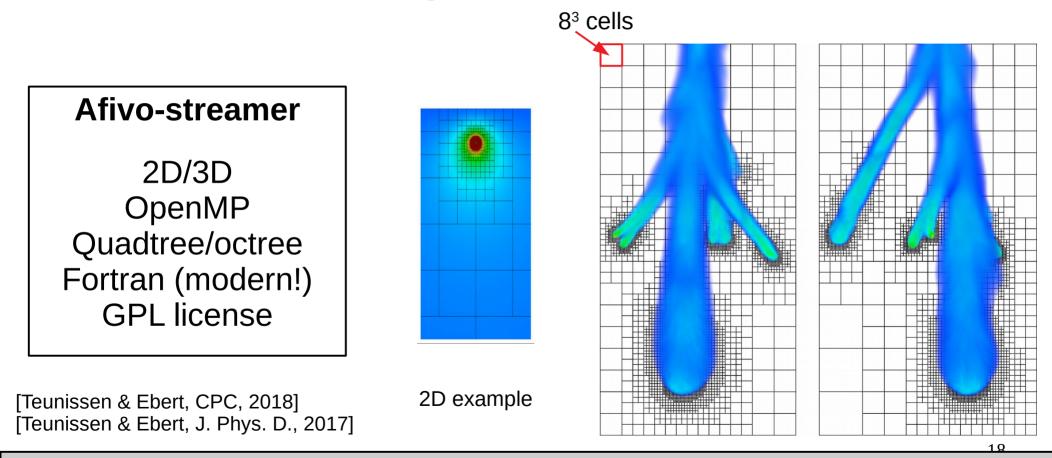
Geometric multigrid



Geometric multigrid

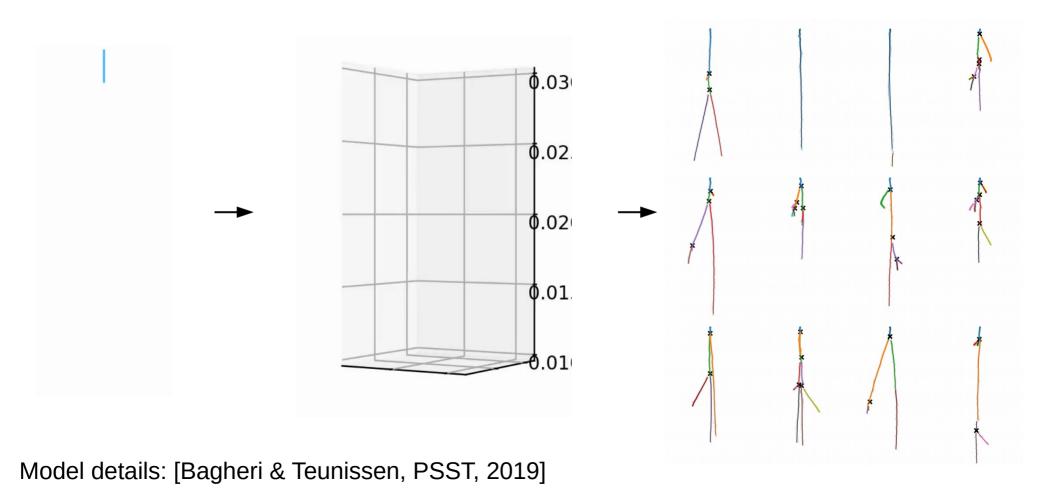


A discharge simulation code



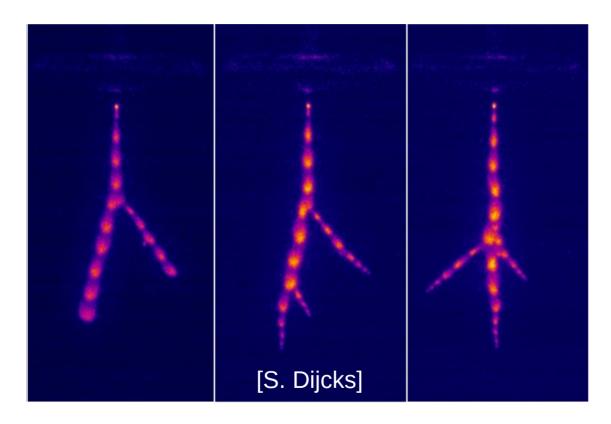
https://gitlab.com/MD-CWI-NL/afivo-streamer

Recent study: branching



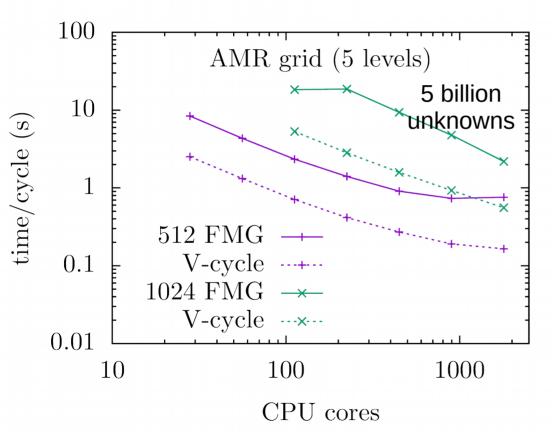
Experiment

Simulation

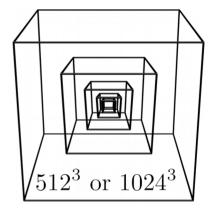


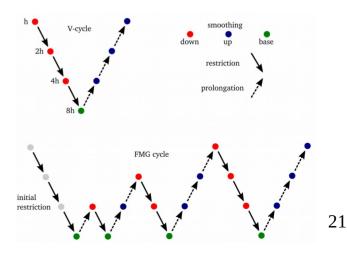


A highly parallel AMR MG solver



[Teunissen & Keppens, CPC, 2019]





http://amrvac.org

- Shock-capturing schemes
- 1D-3D, various geometries
- Includes HD/MHD modules

AMRVAC Meeting @ CWI



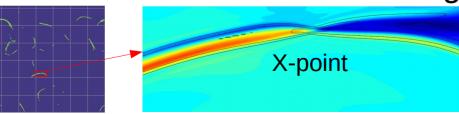
Divergence cleaning
$$\nabla \cdot \nabla \phi = \nabla \cdot \boldsymbol{B}_{old}$$
 $\boldsymbol{B}_{new} = \boldsymbol{B}_{old} - \nabla \phi$



Summary

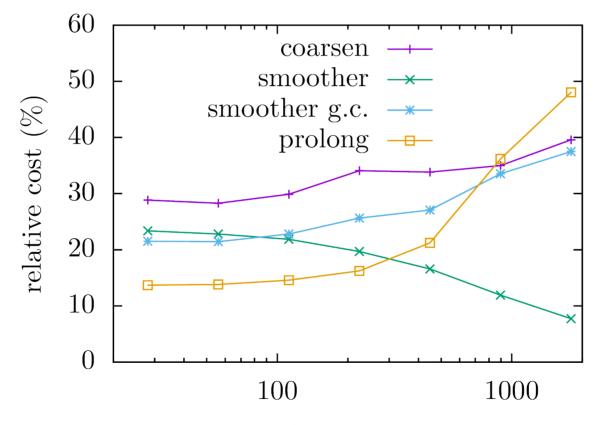
- There exist many types of plasmas
- Plasmas have complex, multiscale behavior
- Computers play a key role in understanding them:
 - From data: machine learning
 - With simulations: physical models
- Modeling often requires AMR and fast elliptic solvers
- Outlook: combine physical models & machine learning

Detecting reconnection in simulations and data (with Andong Hu)



Extra slides

Cost breakdown (1024³ case)



CPU cores

Ghost cells near refinement boundaries

